# Space Weather and Critical Infrastructure in the Australian/ New Zealand Region

**Richard Marshall** 

IPS Radio and Space Services Australian Bureau of Meteorology

# The IPS

#### The lonospheric Prediction Service - Providing HF Radio services and support since late 1940's





# **IPS Radio and Space Services**

# **Providing Space Weather services (including HF Radio) and support since late 1990's**





#### .....Australian BOM's Space Weather Unit since 2007

# **Critical Infrastructure Groups**

- The rise of solar cycle 24 and the growing awareness of impacts of space weather has lead to increased interest from critical infrastructure groups within Australia
- Engagement with Australian Government Trusted Information Sharing Network, includes Energy, Communications, and Transport Sector groups
- Communications sector (CSG) includes, Telstra, Optus (satellite operators), UHF/VHF users (major TV broadcasters), HF users (SES)
- Transport sector (TSG) includes major airlines (Qantas, Virgin), Air Services Australia, shipping and rail
- Energy sector (ESG) includes power network asset owners and operators, gas pipeline owners and operators, mining companies
- All groups operate/own/utilise one or more technologies and systems susceptible to space weather

## **GPS Systems (TSG)**



## **High-End GPS Systems**



# **High-End GPS (TSG – Aviation)**

• **April 2010** - Consultancy to Airservices Australia for GBAS Ionospheric Threat Model Evaluation: Mid-latitude Australian Region





Development of a data-based model for Extreme Space Weather prediction

- May 2012 Proposed Dst-based alerts for anomalous ionospheric gradient conditions at midlatitudes
- Jan-Aug 2012 Development of a data-based model for Extreme Space Weather prediction; Internal report published August 2012; Scientific journal paper in draft
- July 2012 Ionospheric Alert Assessment and Notification Procedure, ASA Protocols







## **Australian Pipeline Network (ESG)**



#### **Australian Pipeline Network**



(Marshall et al., Space Weather, 2009SW000553)

## **Australian Power Network (ESG)**



Increased connectivity over recent years

- Market Competition
- Robustness to demand
- Increased susceptibility to Space weather
- AEMO operates the energy markets and systems in eastern Australia, including PSSWG
- Historically been considered relatively immune to space weather due to mid-latitude location
- Several studies over past few years

#### **New Zealand Power Networks**



**Date/Time Description** 

06-Nov-01, 14:52:00 HWB T4 and ISL SVC tripped. Many South Island transformer NER alarms. HVDC running OK

in balanced mode. HVDC load 216MW with Pole 2 and a half pole in service

06-Nov-01, 14:52:00 Buchholz trip

06-Nov-01, 14:53:00 Red phase caused the tripping. D2 protection flag. Maintenance contractor advised

06-Nov-01, 14:53:00 Buchholz trip. SC advised. HBC to call out contractors

06-Nov-01, 15:08:00 Trip. Alstom called out. SC advised

06-Nov-01, 15:25:00 SI NER alarms reset

06-Nov-01, 15:26:00 T4 tripped at HWB. Ohau unit transformers NER saturation. ISL SVC tripped. Unknown cause.

Requested CLU max VArs at ROX 110 kV, Extra machine started. Third cap switched in at BDE.

Requested extra machines on OHA, OHB, OHC running on TWD to alleviate NER transformer saturation

06-Nov-01, 15:27:00 TKB, CYD, OHA, OHB NER alarms

06-Nov-01, 15:34:00 TOC (Internal transformer fault, explosion vents blown)

06-Nov-01, 15:34:00 HBC advise Red phase unit indicates internal fault. SC told





#### **Power Networks - Australian Region**

#### UPDATE ON ASSESSMENT OF POTENTIAL IMPACT OF CORONAL MASS EJECTION, SOLAR FLARES, GEOMAGNETIC STORMS ON THE NEM POWER SYSTEM



#### **Power Networks - Australian Region**



# **ESG**BULLETIN

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August 2012

#### **MESSAGE FROM THE ESG CHAIRS**

#### Dear ESG Members



SOLAR FLARES

of the sun

power syste

**BEING MONITORED** 

A large solar flare in mid-July has highlighted the importance of monitoring the impact of space storms on the NEM transmission network.

As previously reported in AEMO Energy Update, it was

thern hemisphere but recent studies in New Zealand

and South Africa raised the possibility that Australian power

astructure may be affected by storms on the surface

thought solar flares only affected infrastructure in the

Group Manager Real Time Operations Mark Miller said minor solar storms aren't likely to have an effect on Australia's power infrastructure but large Coronal Mass Ejections (CMEa) have the potential to cause damage. "Coronal Mass Ejections cause geomagnetic storms on earth, which can cause geomagnetically induced current (GIC) on transmission lines. These DC currents create harmonics on the AC network and over-heading on transformers or

reactors. This can lead to transformer and reactor failures and disruption to the network." Mark said.

In recent months AEMO and the Power System Security Work Group (PSSWG), which consists of representatives from Transmission Network Service Providers (TNSPa) together with the Bureau of Meteorology, IPS Radio and Space Services, have been assessing the potential impact of solar storms which tend to peak every 11 years, with the next level of heightened activity occurring over 2012-13. Fortunately the recent solar fare did not impact the AEMO has been working with TNSPs to assess the potential effects of solar storms by installing monitoring equipment on the transmission network. Data is already being collected through the SCADA system in South Australia via Electranet and Queendand via Powerlink. Over the next few months data will become available from other TNSPs.

AEMO also receives warning emails from IPS Radio and Space Services if there is a possibility of a solar flare occurring in the coming days. These warnings are like a weather forecast for solar activity.

The damage caused by solar flares is not to be underestimated. A period of solar storm activity in March 1989 saw virtually the entire Canadian province of Quebec plunged into darkness. In total 21,500 MW of load and generation was lost and it took more than nine hours to fully restore supply.

"The monitoring and mitigation measures we are implementing or considering are prudent and the cost is relatively small when compared to the damage that could be caused by solar storms," Mark said.

For more information about solar storm monitoring see: ACE spacecraft – the red trace in the top box is the north south magnetic field – if it goes below -20nT for an hour, a significant storm will probably result.

http://www.swpc.noaa.gov/ace/MAG\_SWEPAM\_24h.html Magnetic storm Index – http://www.los.gov.au/Geophysical/1/2/4



The core value of the TISN has always been information sharing between critical infrastructure owners and operators. The information improves our understanding of both vulnerabilities and mitigation strategies and provides us with a greater appreciation of one's interdependencies.

The ESG is continually sharing information within the energy sector and with other critical infrastructure sectors through a number of projects and communication avenues. An important project that the ESG is currently working on for the benefit of both ESG and other TISN members is the guide to electricity disruptions that aims to bridge the knowledge gap on how electricity shortages are managed. We hope to distribute this product to other sector groups prior to this summer's storm season and will hopefully have a final draft for members to endorse at the next ESG meeting.

 AEMO update on solar storm activity
 APPEA 2012 conference summary
 Industry access to Australian Government

- Australian Governmen classified information
- Energy Networks 2012

## Ongoing development of SOP's for Space Weather

#### Further monitoring results

- More extensive modelling to identify vulnerable components
- Feedback of findings to SOP's

Australian Energy Market Operator Stakeholder Newsletter 7

## **Extreme Space Weather (ESW) Model**

- Most space weather impacts "of concern" in Australian/NZ region associated with extreme events
- Develop model to forecast extreme events to assist critical infrastructure
- Generalised Linear Model (GLM) techniques not resource intensive
- Event-based analysis (M.Terkildsen)
- Requirements:
  - **LATENCY** 
    - 'Long range' warning (> 12 hours) 
       Based on solar data only (in ASFC)
    - 'Short range' alert (~ 1 hour) ◊ Based on solar data + ACE

#### <u>ACCURACY</u>

- Long range: Optimise to minimise missed events
- Short range: Optimise for forecast accuracy

#### **SIMPLICITY**

- Design for active use in space weather forecast environment

#### **ESW Model Events**



**ESW Model Parameters** 

#### Model covariates (the 'input data')

X-RAY FLARE	SOLAR CYCLE		
Solar flare magnitude	<ul> <li>Solar Cycle Phase</li> </ul>		
Solar flare duration			
LOCATION OF SOLAR ACTIVE REGION	SOLAR WIND / IMF		
Latitude of solar active region	– IMF Bz		
Longitude of solar active region	<ul> <li>Solar wind shock</li> </ul>		
CME CHARACTERISTICS	LJ		
Presence of Halo CME (CME width)			
CME speed			

## ESW Model GLM

#### **Generalised Linear Model**

Response variable (what is being modelled)

$$dstN = \begin{cases} 0 & Dst \ge -50 \\ |Dst + 49| & else \end{cases}$$

Model (a GLM)

$$\ln(\mu_i) = \alpha_0 + \alpha_1 x_{i1} + \alpha_2 x_{i2} + \dots + \alpha_m x_{im}$$

Training data (for fitting model coefficients) 15 years data (1996 – 2010)

Prediction (a binary output)

$$y_{i} = \begin{cases} 0, p_{i}(y \mid \mu) < p_{thresh} \\ 1, p_{i}(y \mid \mu) \ge p_{thresh} \end{cases}$$

 $\begin{array}{l} x_{i1} \dots x_{im} \xrightarrow{} \text{Input variables} \\ a_0 \dots a_m \xrightarrow{} \text{Coefficients (to fit)} \\ p_i(y \mid \mu) \xrightarrow{} \text{Response distribution} \\ \end{array}$   $\begin{array}{l} p_{thresh} \xrightarrow{} \text{Threshold on event} \\ \text{probability used to produce} \\ \text{binary prediction (ESW} \\ \text{event/no-event). Optimised} \\ \text{for required model} \\ \text{performance.} \end{array}$ 

#### **ESW Model Validation**

#### Model validation: Solar data only

#### Optimising for <u>no missed events</u> (false negatives = 0)



#### **ESW Model GLM**

#### Model validation: Solar data + IMF Bz (< -20nT)

#### Optimised for <u>no missed events</u> (false negatives = 0)



## **ESW Model Implementation**

- Operational GUI in ASFC
- Uses Dst-based ESW models as a back-end, proving both binary and probabilistic forecasts for ESW
- Accepts a range of covariates for added flexibility
- Simple text warning message

	Extreme event warnings			Help	
Preliminary HF	Model selection				
HF	Solar data only	⊙ Solar data + IMF Bz			
SWF		Solar data + IMF Bz + solar wind	shock		
Extreme event	Flare/CME observations				
Model 1	Flare date 14-Sep-2012	Latitude n4	(eg, N24)		
Model 2	Flare magnitude x1 (eg, M	5.3) Longitude v13	(eg, W15)		
	Flare duration 1:00 (hh:m	m) Halo CME 🗹 (can	be asymmetric)		
	Solar wind observations (when CM	E hits ACE)			
	Southward IMF Bz Solar wi	ind shock 🗌			
	Duri Madal	Send warning?			
	EXTREME EVENT			Recent warning time	
		Satellitor	Send Override	Recall	
	Model Output	Satemies			
	Model Output	Power Grids (AEMO)	Send Override	Recall	
	Model Output Reset	Power Grids (AEMO)	Send Override Send Override	Recall	
	Model Output Reset	Power Grids (AEMO) GPS	Send Override Send Override	Recall Recall	
	Model Output Reset	Power Grids (AEMO) GPS Aviation	Send Override Send Override Send Override	Recall Recall Recall	
	Model Output Reset	Power Grids (AEMO) GPS Avlation	Send Override Send Override Send Override	Recall Recall Recall	



## **Future Developments**

- ACE solar wind parameters
- CME symmetry parameter / CME "mass"

ters ter / CME "mass" (e.g. Kim et al 2008, 2010)

Active region magnetic characteristics (proxy for IMF Bz events?)



- CME travel time (flare-shock interval) to replace CME plane-of-sky speed (a poor proxy for true CME speed)
- Type II / Type IV radio bursts
- Direct modelling of ESW parameters as response
  - GIC index
  - Ionospheric gradient index