



to help Protect Satellites on Orbit

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- Space is strategically important for Europe
 - Industry, GMES, Galileo,....
- Space assets are vulnerable to high energy particles
- Vulnerability increasing new technology
- Risk changes with solar cycle
- Effects of another Carrington event?
 - Estimated \$30 bn [Odenwald and Green, 2007]



Galileo - Courtesy of ESA



Radiation Damage to Satellites



- Solar energetic particles
 - Single event upsets
 - Degrades solar array power
- High energy electrons (Rad. belts)
 - Internal charging ESD
- Low energy electrons
 - Surface charging ESD
- Total ionizing dose
- Micrometeoroids and debris ESD





Satellite Anomalies – When SW Conditions Disturbed

- 20th Jan 1994
 - Intelsat 4, Anik E1 and Anik E2
 - Intelsat 4 and Anik E1 were recovered in a few hours
 - Anik E2 Loss of service for 6 months
- 11th January 1997
 - Telstar 401 Total loss Insurance payout \$132m
- 19th May 1998
 - Galaxy IV Total loss Insurance payout \$165m
- 23rd Oct to 6th Nov 2003
 - 47 satellites reported malfunctions
 - Midori 2 Total loss US\$640m scientific satellite
- 5th Apr 2010
 - Galaxy 15 Loss of service for 8 months drifted around GEO risk of collision
- 7th March 2012,
 - Sky Terra 1 and Spaceway 3 Safe mode, loss of service for hours days





Sunspot Cycle – Geomagnetic Activity



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Earth's Radiation Belts

The Earth's Electron Radiation Belts International Geo Gallileo 4 Space Station GPS $\mathbf{2}$ Earth Radii -2 -4 -2 -8 -4 0 2 4 6 Earth Radii





Magnetic Storms – Radiation Belts



Forecasting Concept



- It takes ~ 40-60 minutes for the solar wind to flow from the ACE satellite to the Earth
- Access ACE satellite data in real time and use it to drive our forecasting models
- Use a forecast of Kp index from Swedish Inst. Sp. Phys. (Lund) and BGS (UK) and data from Europe, USA and Japan
- We use physical models
 - Like weather forecasting



Physical Models Include Wave-Particle Interactions



Radial Diffusion: Transport of Electrons Across the Magnetic Field



- Fast solar wind drives ultra-low frequency (ULF) waves
 - Oscillations in the Earth's magnetic field
- ULF waves enhance electron transport across the magnetic field
 - Radial diffusion



Testing the BAS Radiation Belt Model

 J_{\perp} (cm⁻²s⁻¹sr⁻¹keV⁻¹)

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BAS Model with Radial Diffusion only





SPACECAST – Forecast >800 keV electrons







SPACECAST – Forecast >800 keV electrons



SPACECAST > 2 MeV Electrons



Risk of Satellite Charging - ESD

Model results converted into a risk index based on previous satellite anomalies

Last 24 hours Forecast 10⁵ Geostationary Orbit (GOES) High risk >2 MeV Fluence (cm⁻² sr⁻¹) 10⁸ Low risk ♦ GOES 10 10' 2000 ឧ 1500 (LL) 6 4 ¥ 1000 Ч 500 0 08:15 18:15 23:15 04:15 09:15 14:15 13:1508 Mar 2012 Plot created on Thu Sep 13 09:39:28 2012 UTC Time 09 Mar 2012 SPACECAST Project

Geosynchronous Orbit

British Antarctic Survey

GNSS/Galileo Orbit



- Risk depends on satellite design
- Needs close collaboration with satellite operators and designers



Benefits of Physical Models

- Forecast what is likely to happen enables mitigation
- Reconstruct what happened in the past identify the cause of satellite anomalies
- Construct data where there are little/no observations GNSS orbits
- Calculate extreme conditions based on physical principles
- Calculate number of particles precipitating into the atmosphere effects on low altitude satellites, ionization and GPS signals





Conclusions

- SPACECAST makes real time forecasts of the radiation belts for satellite operators
- Forecast for 3 hours, updated every hour, and translated into a risk index
- Unique features
 - Physical models, that include wave-particle interactions
 - Forecast for the whole radiation belts including GNSS/Galileo orbits
 - European led with USA and Japan
- Forecasts can be improved by
 - Coupling the solar wind/magnetopause to the radiation belts
 - Including low energy electrons surface charging
- Options to model extreme events and orbits where there is little/no data





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